

# TROPOMI is ready for launch!

## Pre-flight performance and calibration measurements

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# Outline of the presentation

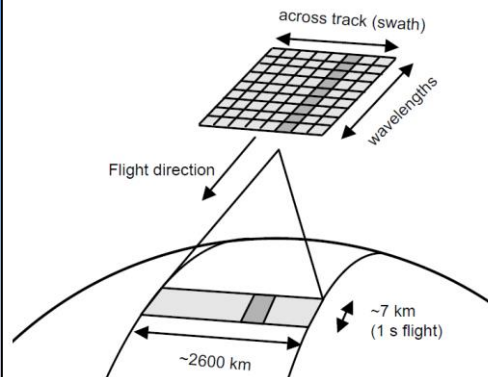
- ❑ Sentinel-5 precursor – mission and status
- ❑ The TROPOMI instrument
- ❑ TROPOMI instrument performance and calibration
  - Calibration in progress
  - A few selected calibration results
    - Intra-band co-registration
    - Deriving the instrument spectral response function
    - Deriving the instrument straylight performance
- ❑ Some concluding remarks

# Sentinel-5 precursor – mission and status

- ❑ The ESA Sentinel-5 Precursor (S-5p) is a precursor mission that focuses on global observations of the atmospheric composition for air quality and climate.
- ❑ Sentinel-5p bridges the data gap between SCIAMACHY / OMI (2002 / 2004) and Sentinel-5 (2021), with improved sensitivity and smaller ground sampling distance
- ❑ The TROPOspheric Monitoring Instrument (TROPOMI) is the payload of the S-5p mission and is jointly developed by The Netherlands and ESA. Instrument prime is Airbus DS Netherlands.
- ❑ On-ground calibration testing of TROPOMI finished on May 2015.
- ❑ The planned launch date for S-5p is end of 2016 with a 7 year design lifetime.

## The instrument:

- UV-VIS-NIR-SWIR nadir viewing pushbroom grating spectrometer.
- Spectral range: 270-500, 675-775, 2305-2385 nm
- Spectral Resolution: 0.25-1.1 nm
- Spatial Resolution (nadir): 7x7km<sup>2</sup>
- Global daily coverage at 13:30 local solar time.

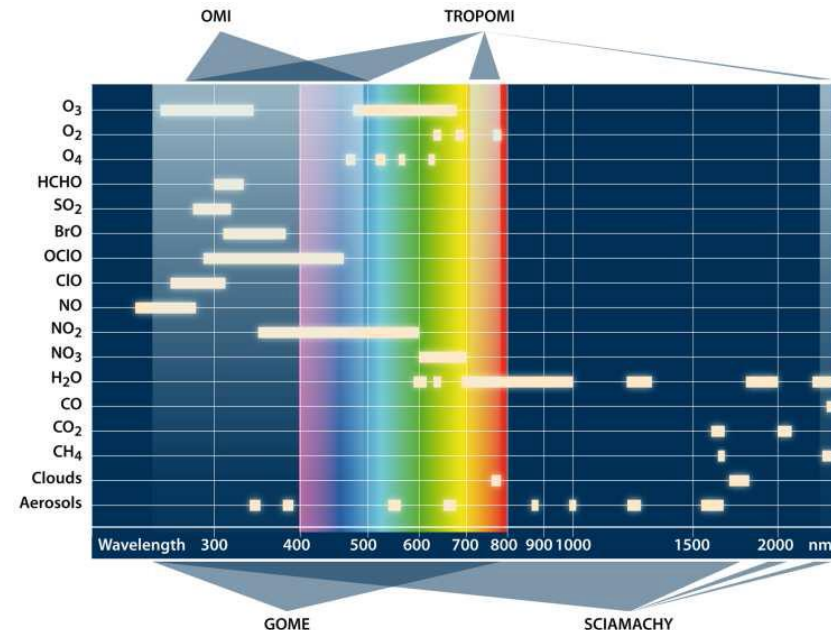
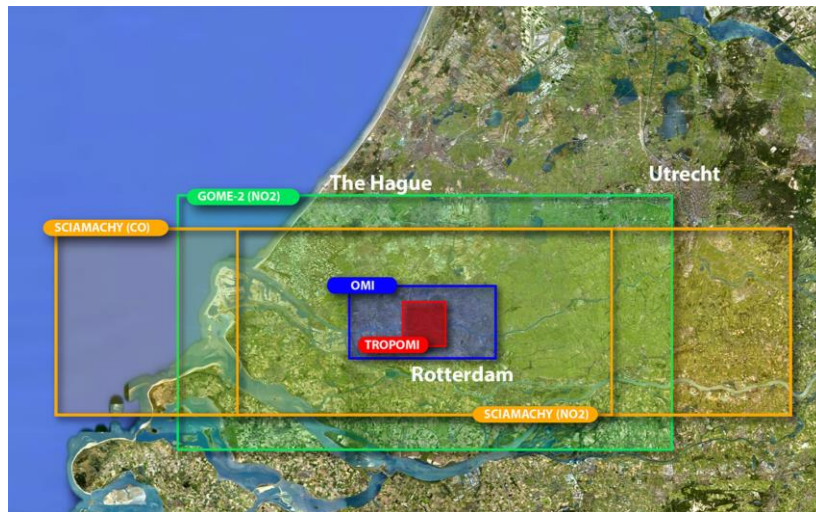


## The science products

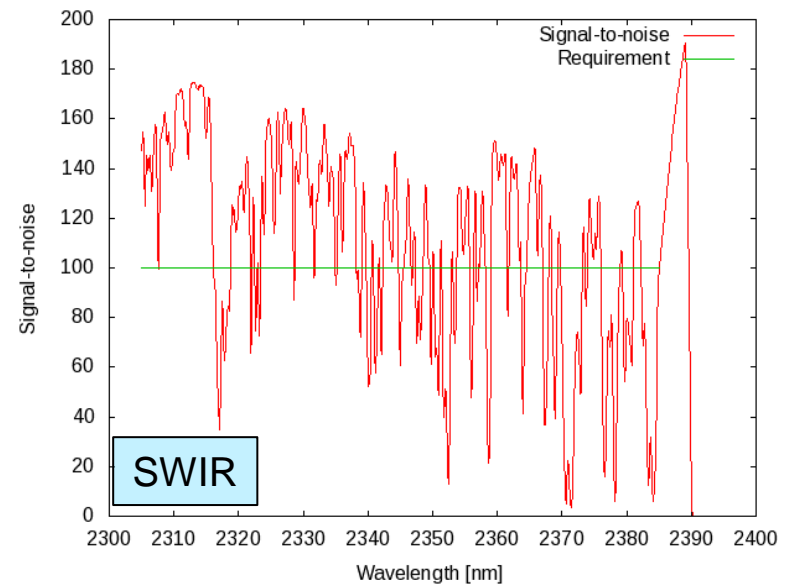
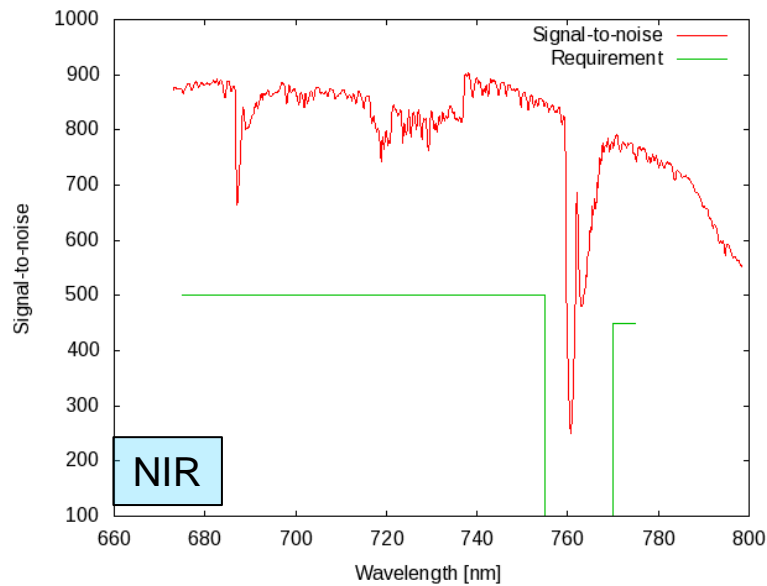
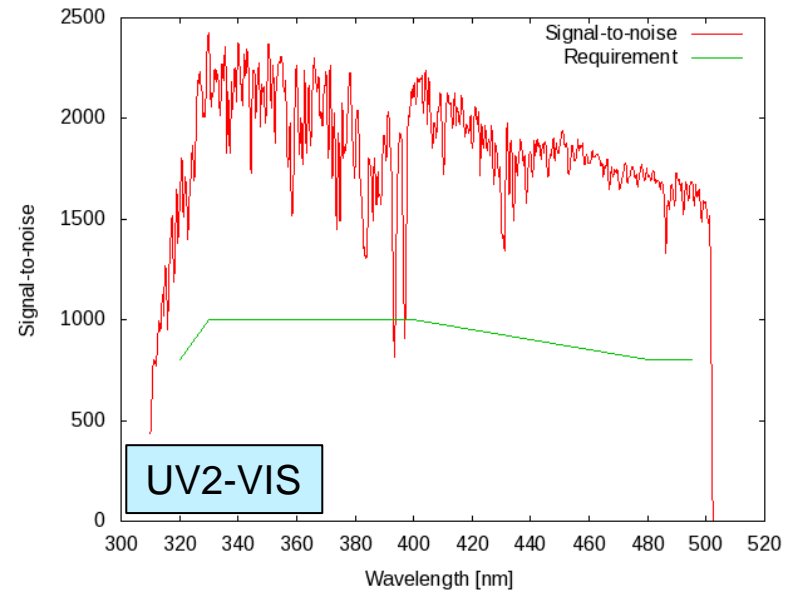
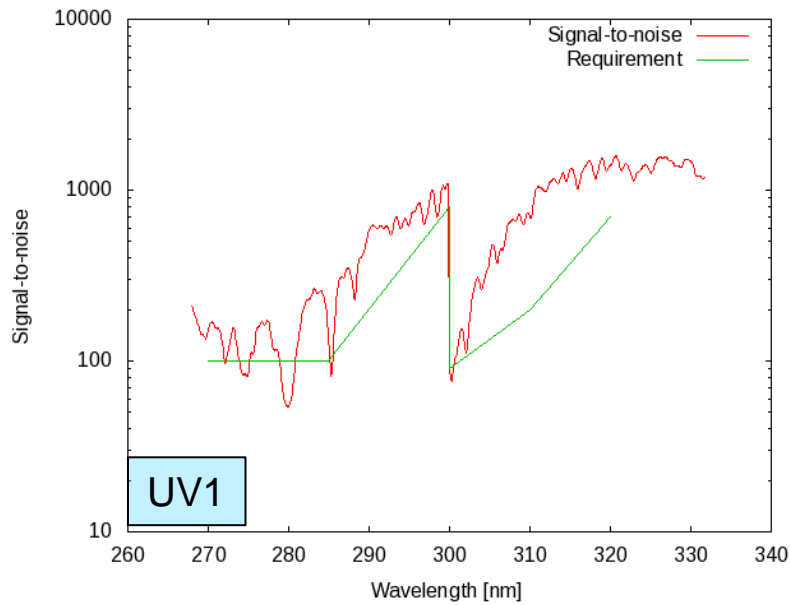
- Total column: O<sub>3</sub>, NO<sub>2</sub>, CO, SO<sub>2</sub>, CH<sub>4</sub>, CH<sub>2</sub>O, H<sub>2</sub>O, BrO
- Tropospheric column: O<sub>3</sub>, NO<sub>2</sub>
- O<sub>3</sub> profile
- Aerosol: absorbing index, type, optical depth

# The TROPOMI Instrument performance parameters

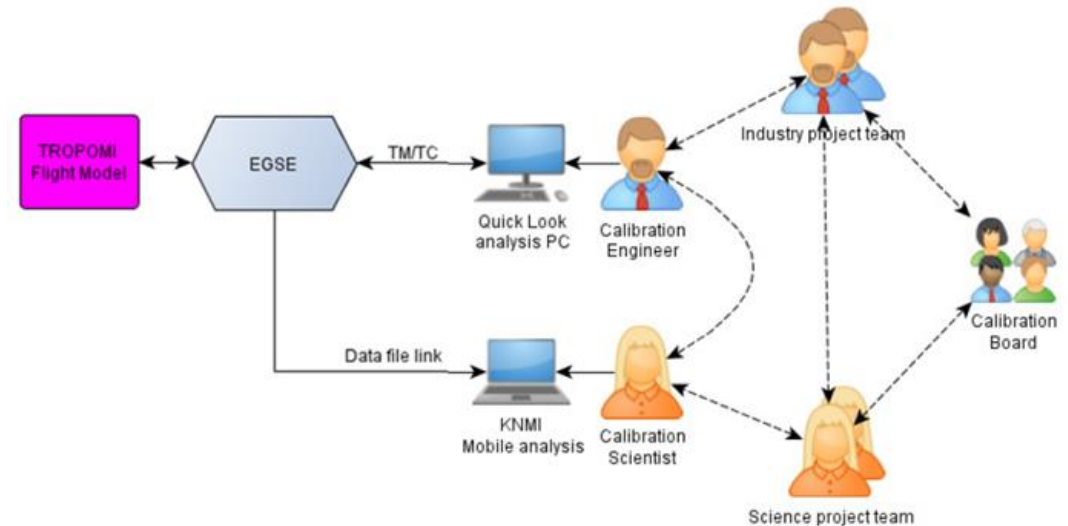
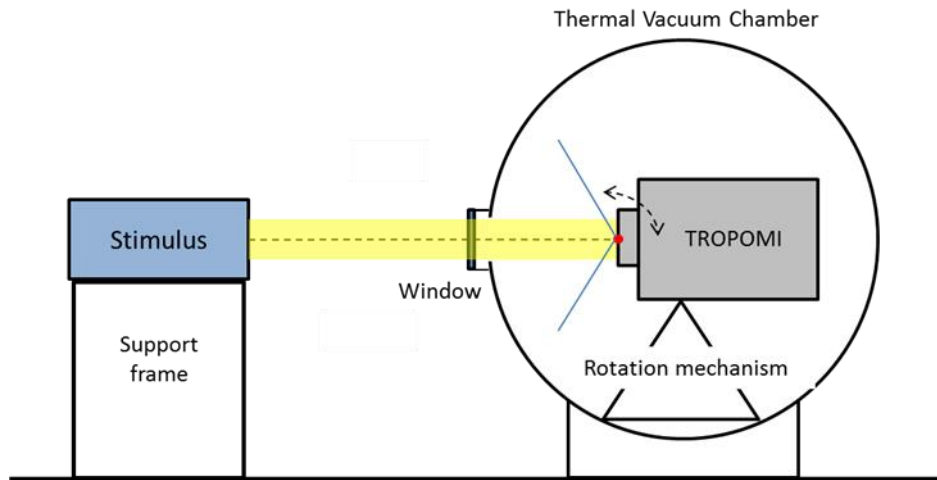
Spectrometer	UV		UVIS		NIR		SWIR
Band ID	1	2	3	4	5	6	7
Spectral range [nm]	270-300	300-320	310-405	405-495	675-725	725-775	2305-2385
Spectral resolution [nm]	0.5	0.5	0.55	0.55	0.5	0.5	0.25
Spectral sampling [nm]	0.06	0.06	0.2	0.2	0.1	0.1	<0.1
Spatial sampling [km <sup>2</sup> ]	21x28	7x7	7x7	7x7	7x7	7x1.8	7x7
Signal-to-noise (required)	100	100-1000	1000-1500	1500	500	100-500	100-120



# TROPOMI SNR for dark scenes (2 - 5 % albedo)



# Instrument H/W is only useful after proper calibration



## Efficiently organized performance / calibration measurement campaign

- ❑ Only 125 days to get all and the best possible necessary Calibration Key Data

## Cooperation

- ❑ Airbus DS: measurement campaign and management
- ❑ KNMI / SRON (PI): measurement definition

**Happily  
cooperating  
teams:  
key to succes!**



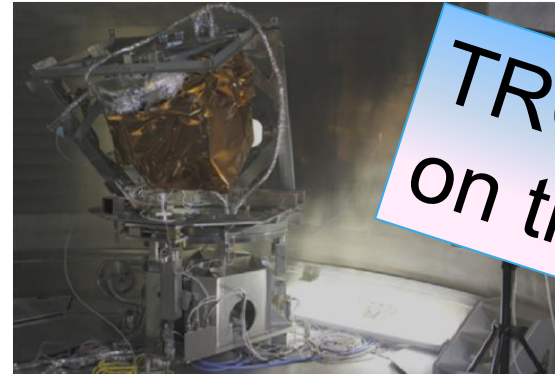
# Calibration in progress

- ❑ Automated measurements in 6 m Ø vacuum vessel at CSL in Liege, Belgium

Preparing TROPOMI



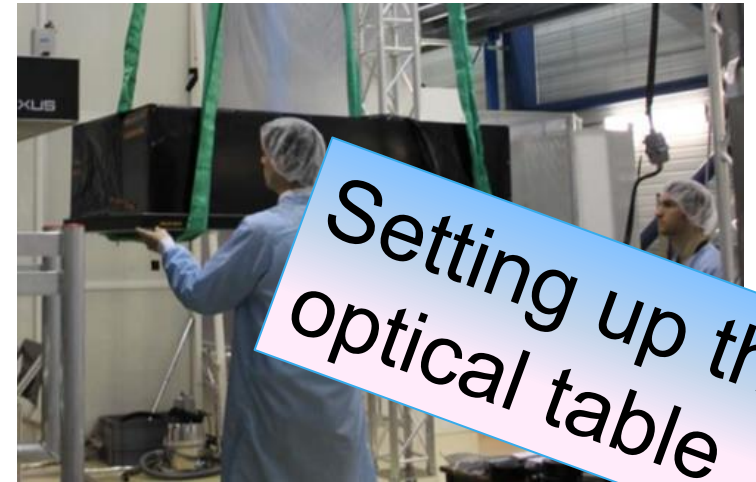
TROPOMI on the cradle



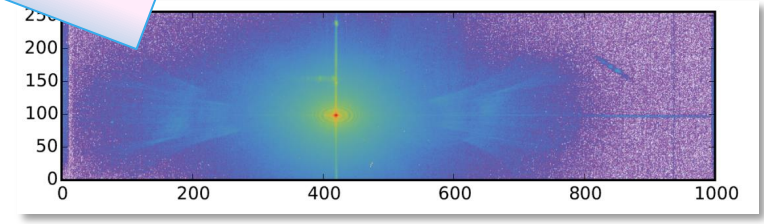
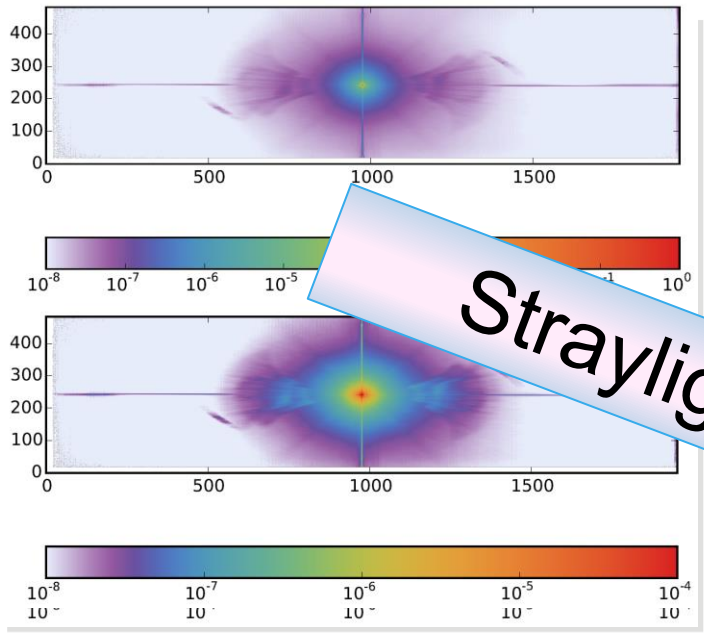
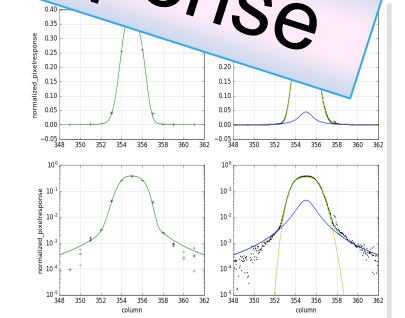
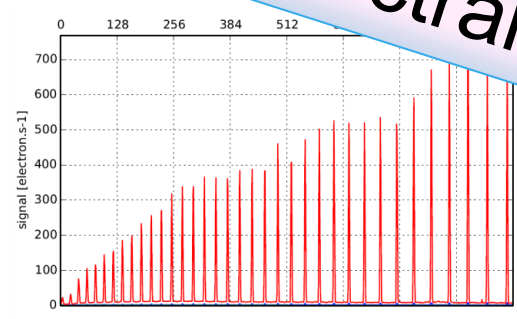
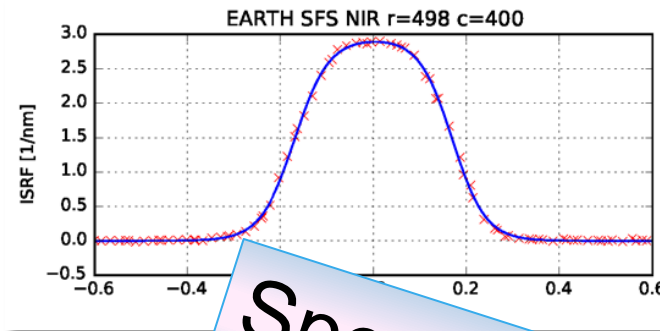
TROPOMI Wrapped in MLI



Setting up the optical table

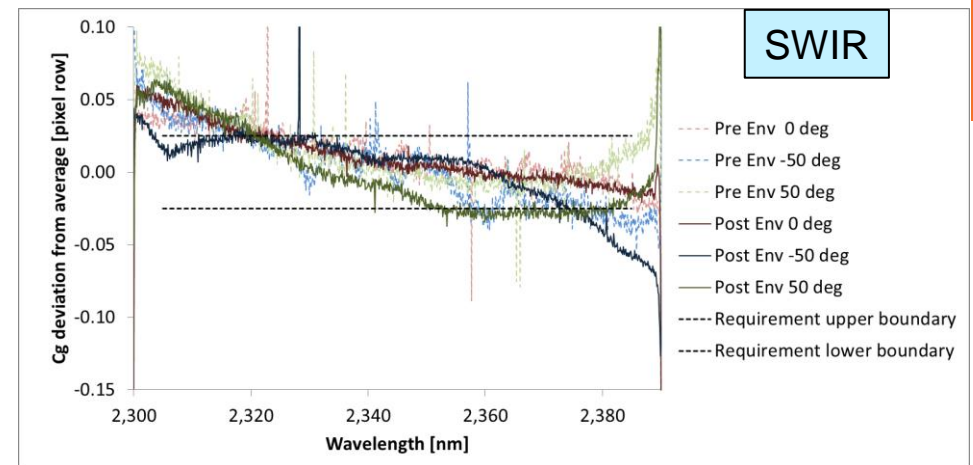
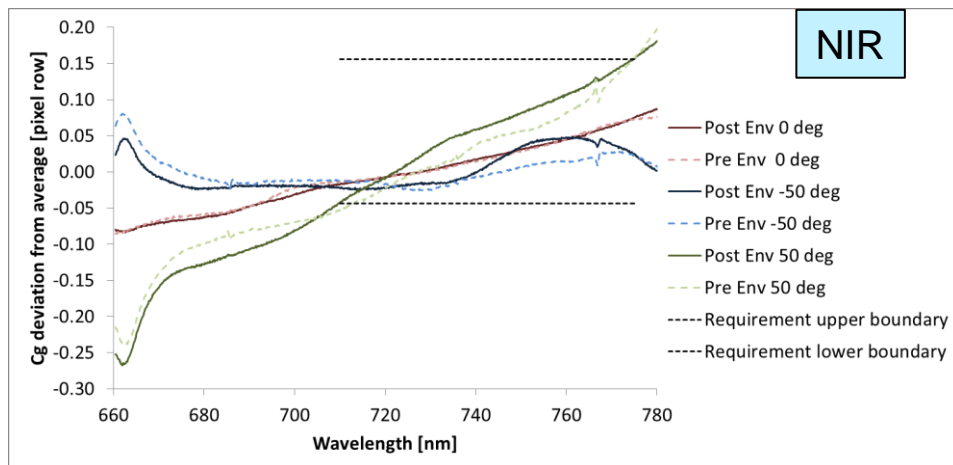
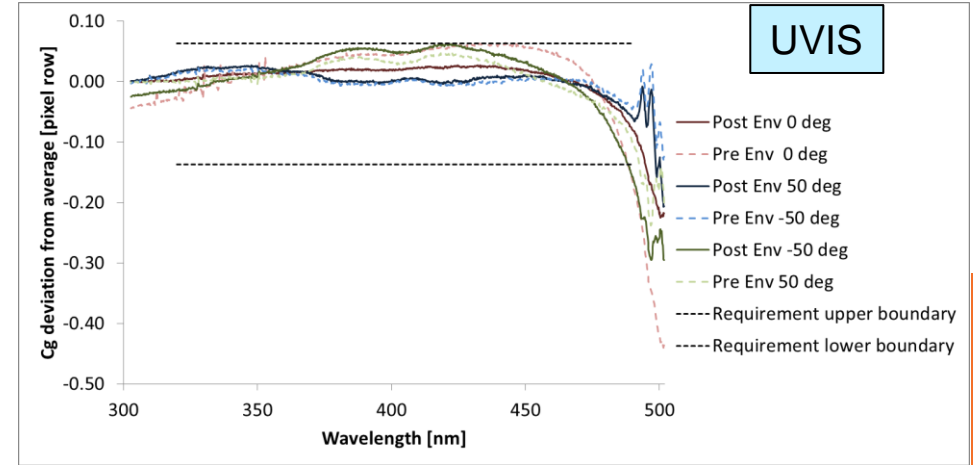
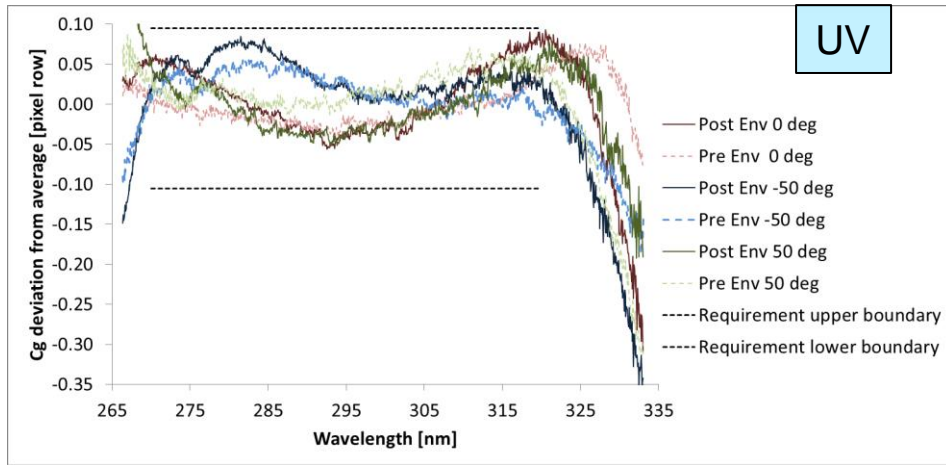


# Some results of the measurements campaign





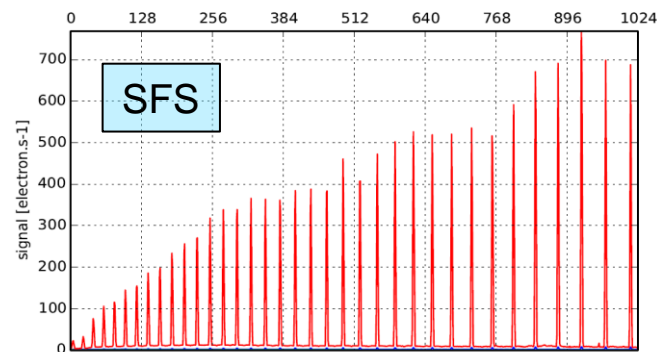
# Intra-band co-registration: within specification



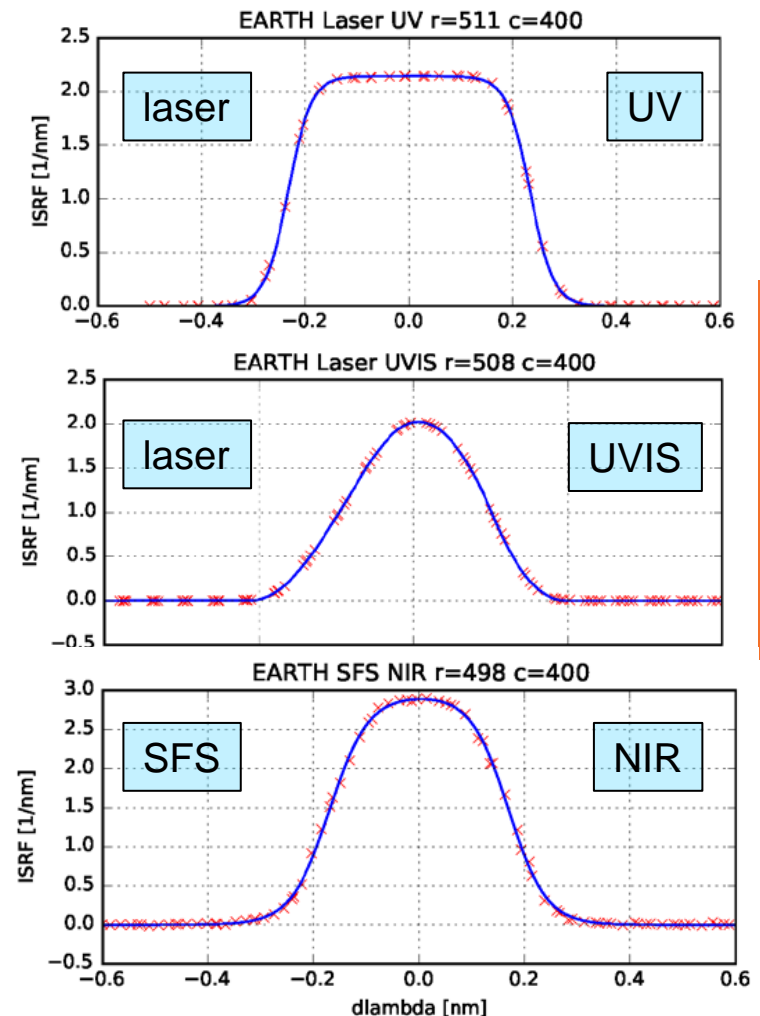
# Deriving the UVN ISRF

## Method:

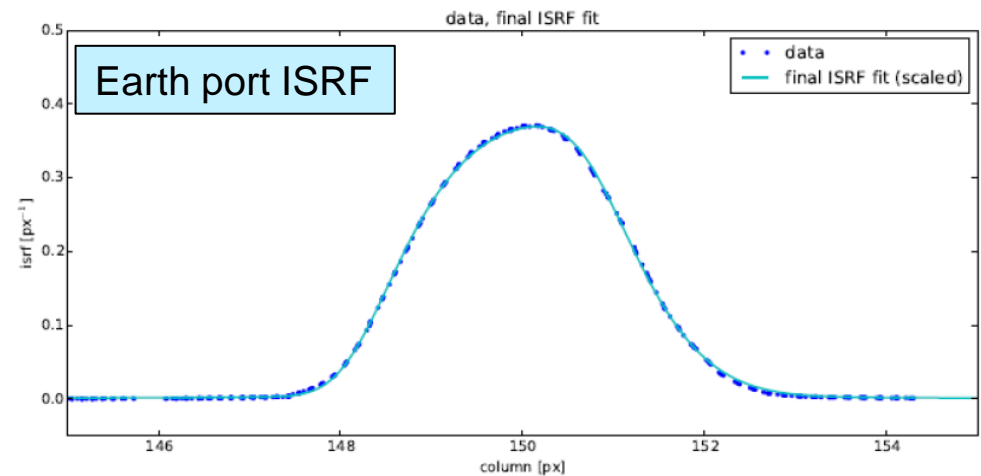
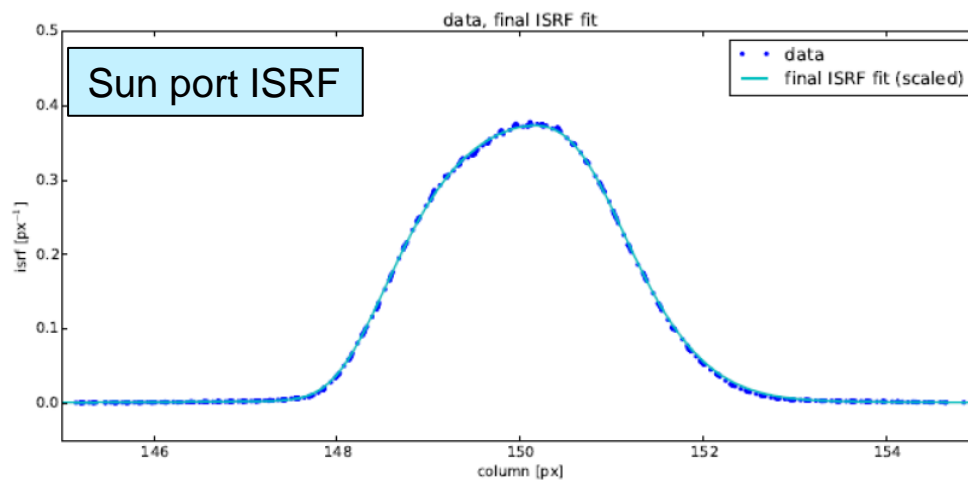
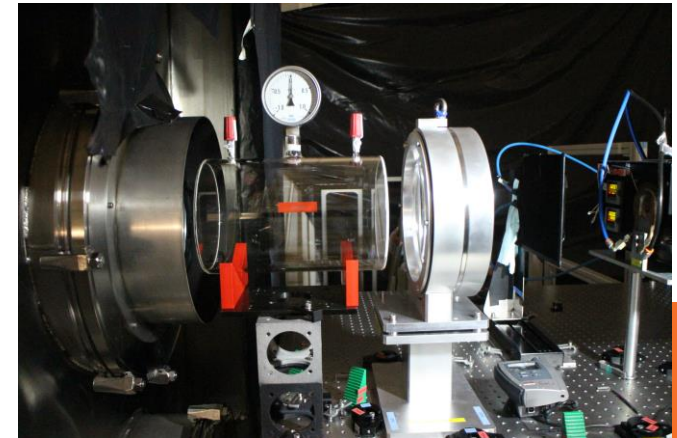
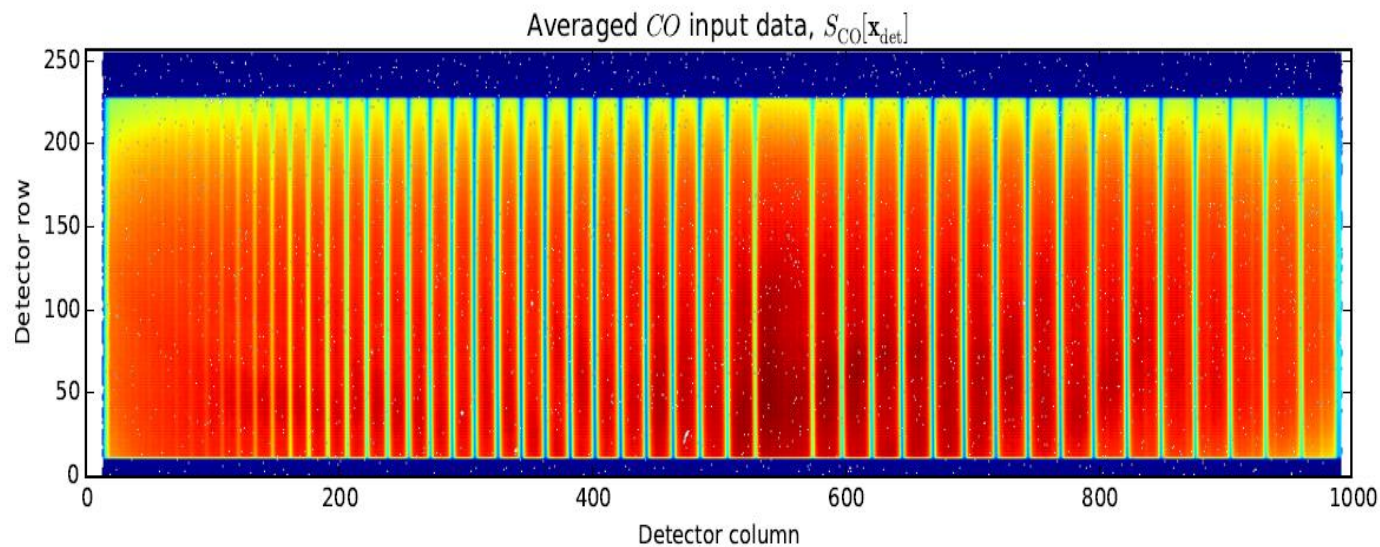
- ❑ A narrow spectral line is tuned in small steps.
- ❑ Line source may be either a laser or a dedicated slit function stimulus (SFS)
- ❑ During tuning, the signal of each individual pixel is followed as the laser scans over this pixel. This is the Instrument Spectral Response Function (ISRF) for each pixel.
- ❑ ISRF is function fitted iteratively
- ❑ Data selection and 2D fitting is needed to remove laser features.



## Results:



# CO gas cell (spectral assignment) and ISRF SWIR band

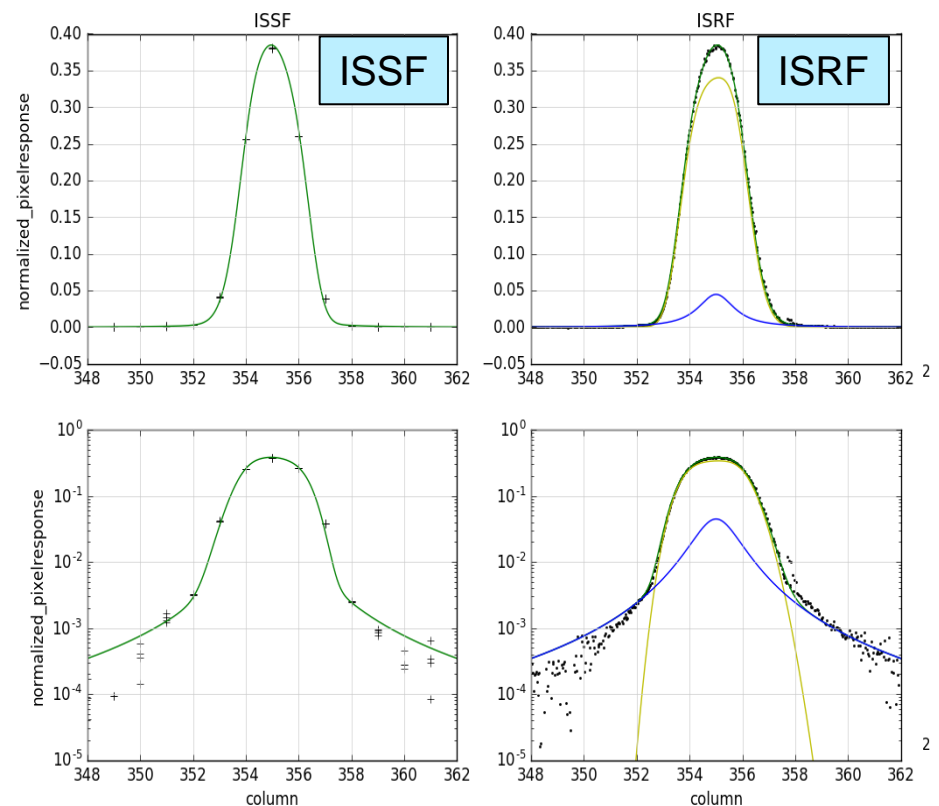


# Deriving the SWIR ISSF and ISRF

## Method:

- ❑ A tunable SWIR laser scans the complete wavelength in small steps.
- ❑ The Instrument Spectral Spread Function (ISSF) is measured for illuminated pixels.
- ❑ During laser tuning, the signal of each individual pixel is followed as the laser scans over this pixel. This is the Instrument Spectral Response Function (ISRF) for each pixel.
- ❑ ISSF and ISRF are function fitted iteratively
- ❑ Data selection and 2D fitting is needed to remove laser features.

## Results:



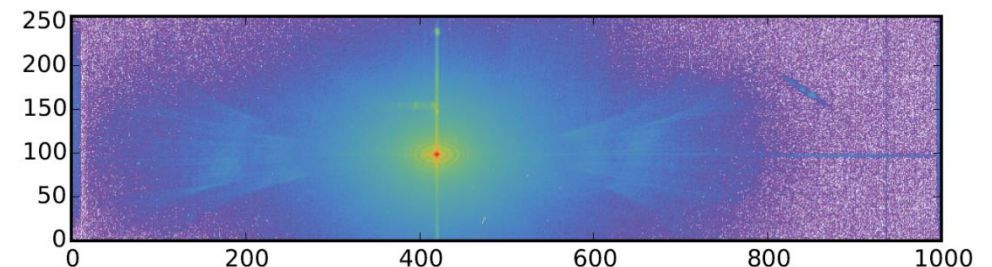
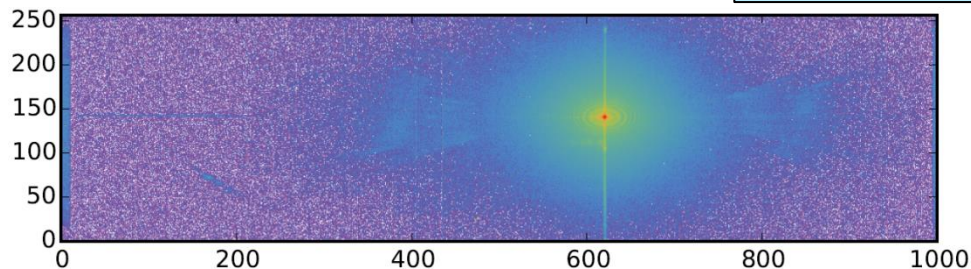


# Stray light SWIR: laser measurements

## Method:

- ❑ Use a tunable monochromatic source (laser)
- ❑ Measure at  $\pm 100$  spatial positions in the FoV and at  $\pm 100$  spectral positions in the band. Total  $\pm 10,000$  point measurements.
- ❑ Measure laser peak with very short exposure time (0.2 ms) and stray light with longer integration time (4, 100, 2000 ms).
- ❑ Combine unsaturated data: 10,000 'superframes' with  $10^8$  dynamic range, see next page.
- ❑ Superframes look very similar! This allows to:
  - Derive 'stray-light kernel' based on 'average' superframe.
  - Deconvolve measurements with kernel.

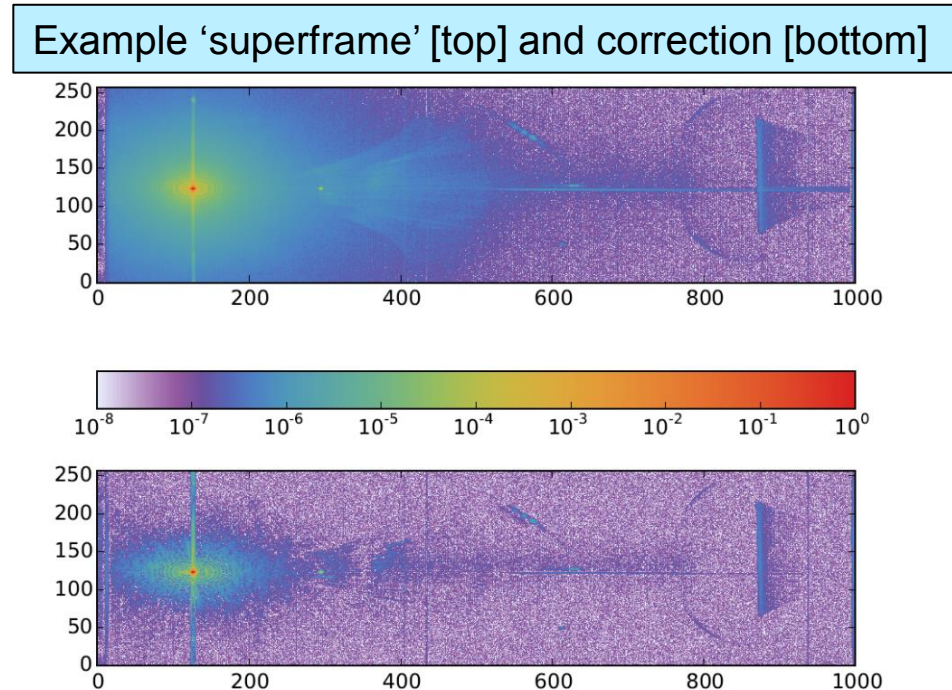
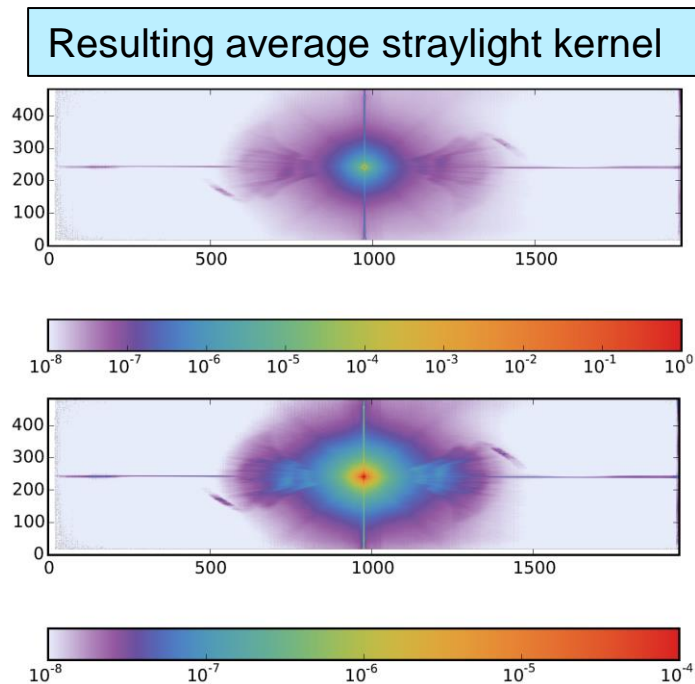
Example of two 'superframes'



# SWIR Stray light kernel and correction

## Results:

- ❑ In-field and in-band SWIR stray light measured with high dynamic range (8 orders) using a tunable laser and  $10^4$  difference in exposure times.
- ❑ Stray light caused by scatter of optical surfaces, pre-spectrometer optics, reflections, grating.
- ❑ On Level-1b stray light out-of-spec by factor of  $\pm 4$ .



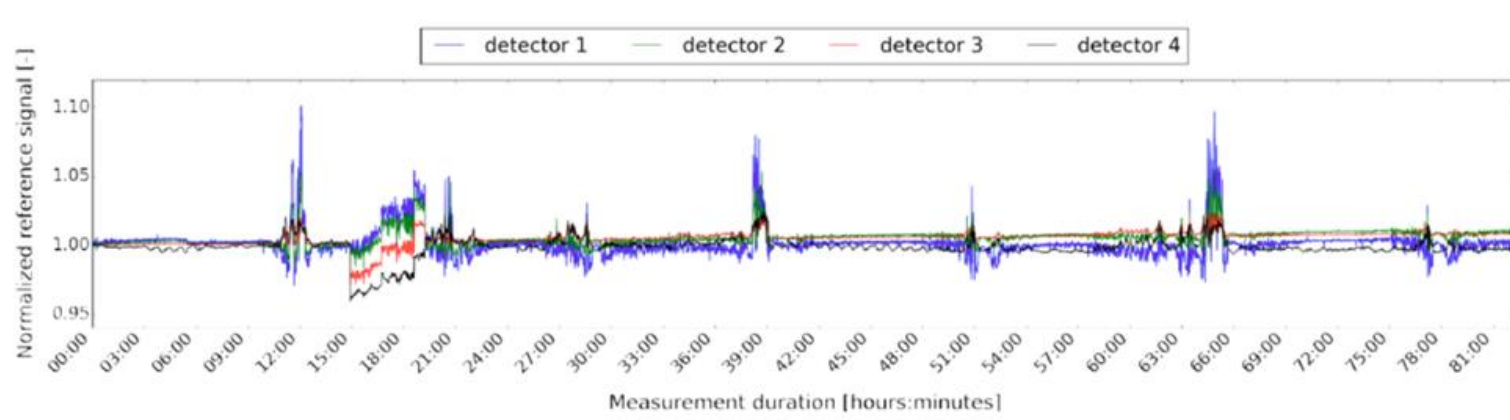
# Successes and drawbacks

## Successes:

- ❑ No functional failures of the TROPOMI instrument
- ❑ Efficient use of versatile light sources
- ❑ Measurement efficiency: (measurement time / total time): 85 %
- ❑ Total 80 % of the planned data acquired

## Drawbacks:

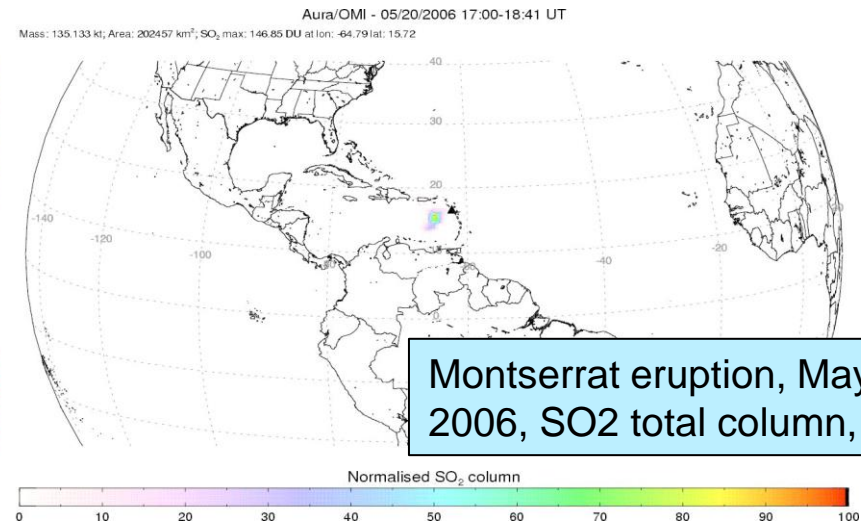
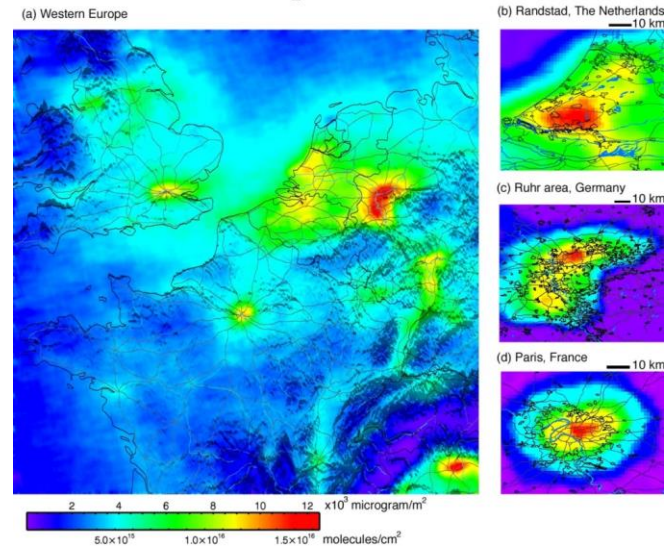
- ❑ Instability of Xe lamp
- ❑ Failure of tunable laser



# TROPOMI is well calibrated, ready for launch and use

- ❑ Launch planned October 2016, from Plesetsk in Russia
- ❑ TROPOMI trace gas / aerosol / cloud data:
  - Climate research
  - Air quality data (independent)
  - Medium range weather forecasts
- ❑ With the imminent launch of TROPOMI, the near future for greenhouse gas measurements from space looks bright.

OMI Tropospheric NO<sub>2</sub>, Dec 2004 until Nov 2005



Montserrat eruption, May 2006, SO<sub>2</sub> total column, OMI