

# **Towards a remote sensing solution to quantify N<sub>2</sub>O emissions by integrating shortwave and longwave infrared bands**

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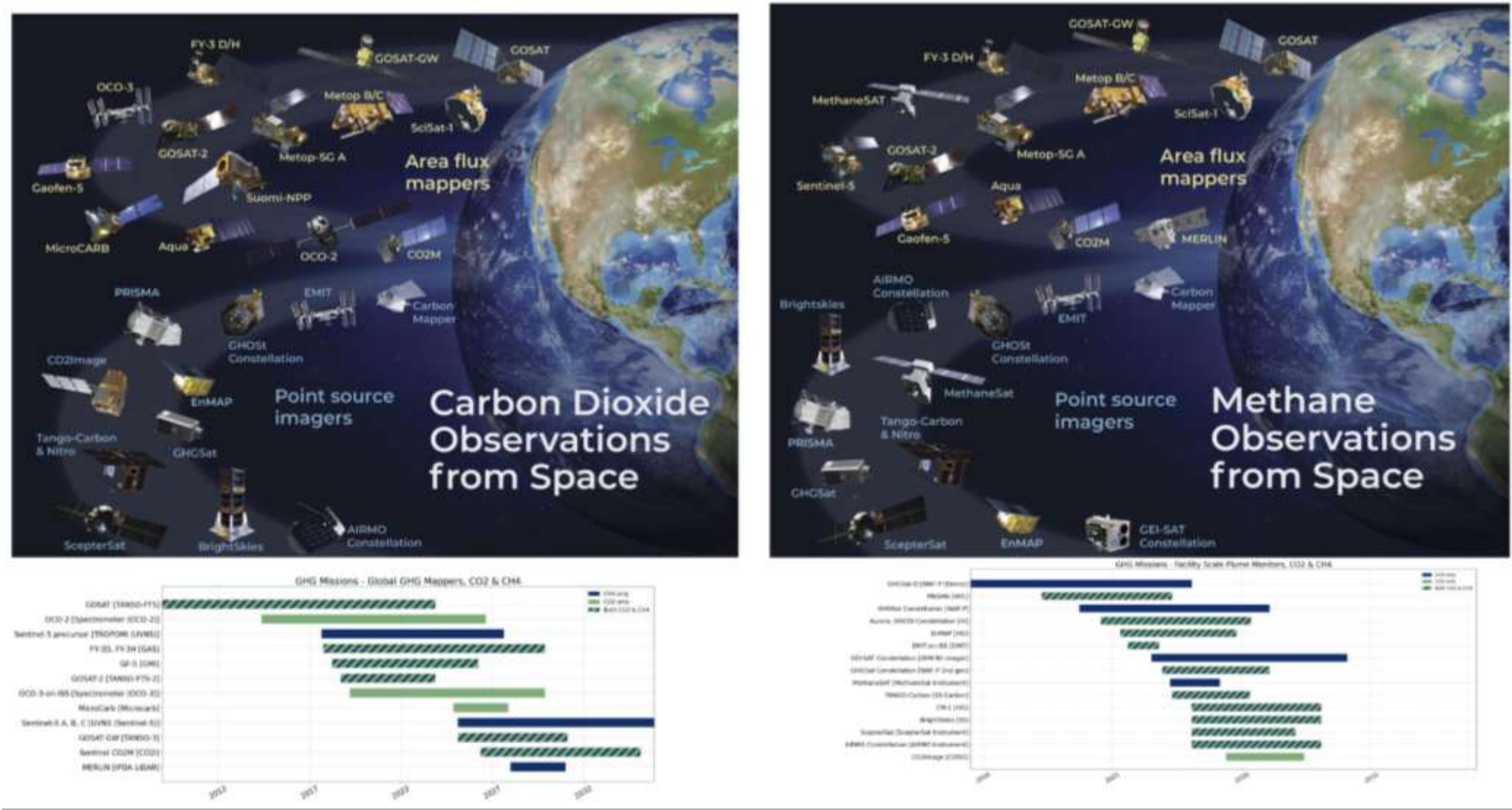
Radiative transfer modelling: Christopher Chan Miller, Robert Spurr, Karen Cady-Pereira

Instrument design: Thomas U. Kampe, Nathan Leisso

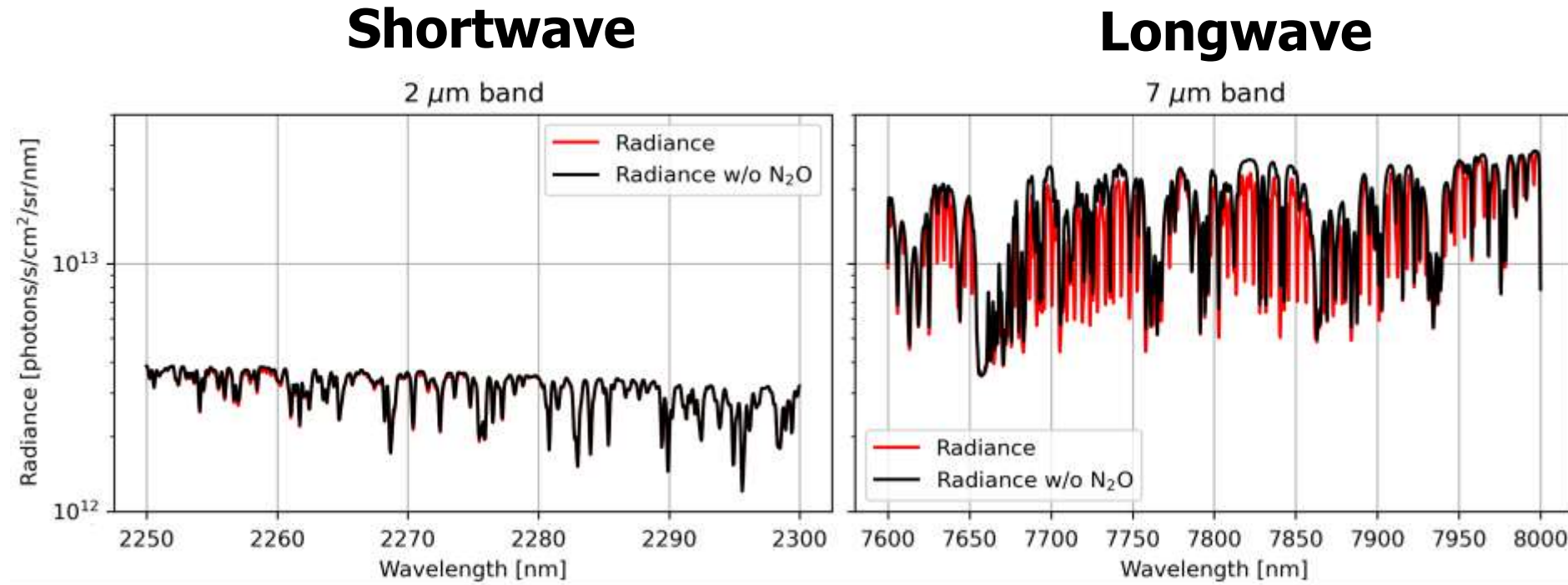
N<sub>2</sub>O flux data: Wendy H. Yang, Emily R. Stuchiner, Eddy Will



# N<sub>2</sub>O: a key greenhouse gas overlooked from space



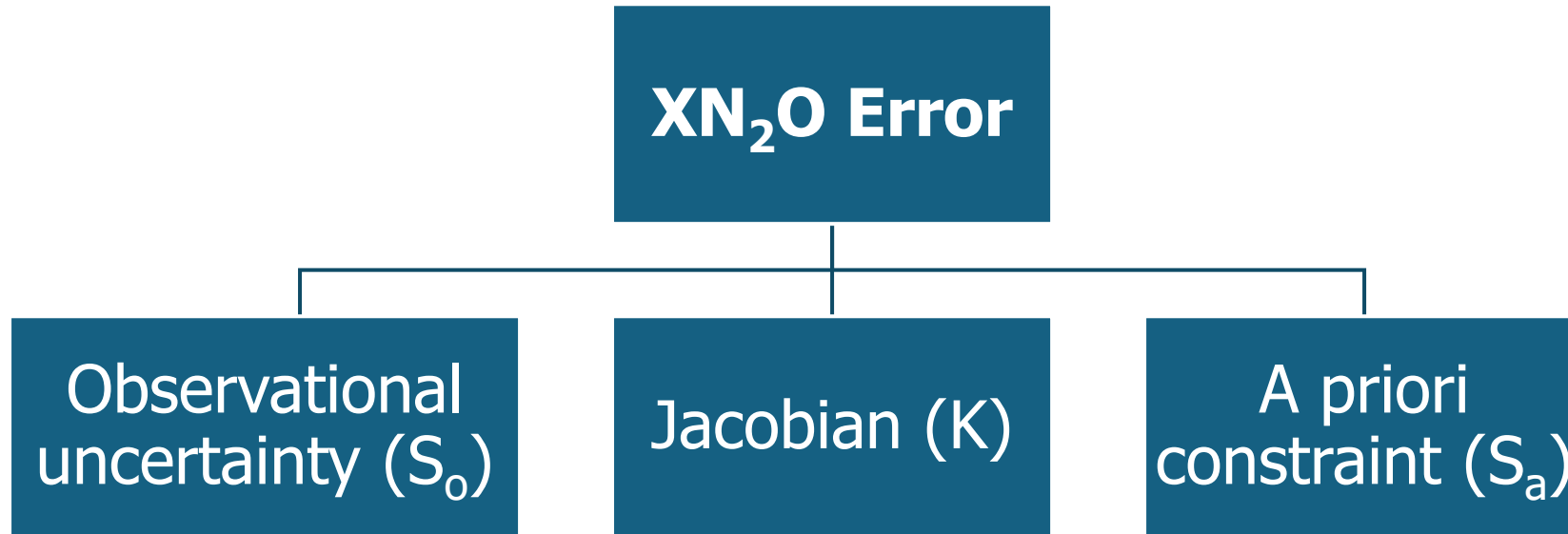
# The challenge of N<sub>2</sub>O remote sensing



- 2  $\mu\text{m}$ : consistent vertical sensitivity but very weak N<sub>2</sub>O absorption
- 7  $\mu\text{m}$ : strong absorption and radiance level but weak near-surface sensitivity
- Integrate 2  $\mu\text{m}$  and 7  $\mu\text{m}$  to combine the strengths of short and longwave bands

# How precisely can we observe $\text{XN}_2\text{O}$ ?

- $\text{XN}_2\text{O}$ : Column-integrated mixing ratio of  $\text{N}_2\text{O}$   $(\text{XN}_2\text{O} = \frac{\Omega_{\text{N}_2\text{O}}}{\Omega_{\text{A}}})$



$\mathbf{S}_m$ : measurement error covariance matrix

Measurement error of  $\text{XN}_2\text{O}$

$$\mathbf{S}_m = \mathbf{G}\mathbf{S}_o\mathbf{G}^T$$

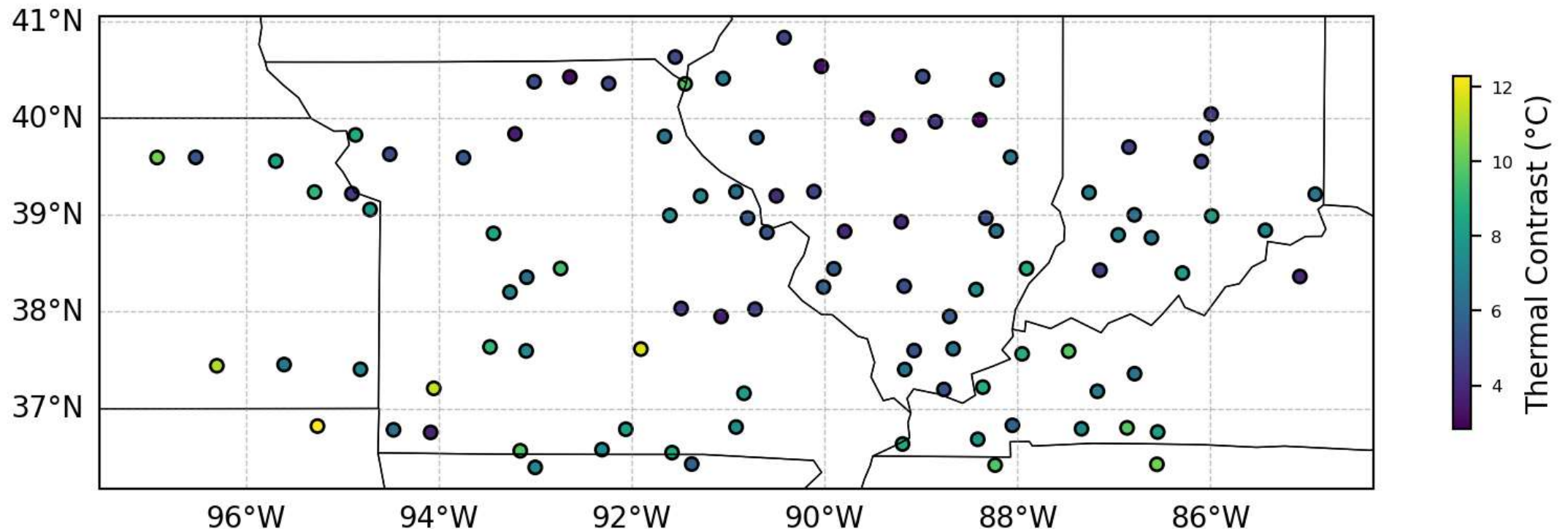
$$\sigma_{\text{XN}_2\text{O}(m)} = \mathbf{h}^T \mathbf{S}_m \mathbf{h}$$

$$\mathbf{G} = \mathbf{S}_a \mathbf{K}^T (\mathbf{K} \mathbf{S}_a \mathbf{K}^T + \mathbf{S}_o)^{-1}$$

$\mathbf{h}$ : pressure weighing function

# Atmospheric and surface states

- CrIS Level 2 files for  $\text{N}_2\text{O}$ ,  $\text{CH}_4$ ,  $\text{H}_2\text{O}$ , and temperature from the JPL MUSES algorithm
- Atmospheric profiles and surface conditions extracted from 100 CrIS pixels
- Area: US Midwest (Part of corn belt)
- Observation date: 23 August 2023





# Observational uncertainty

| Parameter [unit]                              | 2 μm band | 7 μm band |
|-----------------------------------------------|-----------|-----------|
| Wavelength range [nm]                         | 2240–2300 | 7600–8000 |
| Spectral sampling [nm]                        | 0.0575    | 0.25      |
| Slit width [× spectral sampling] <sup>a</sup> | 3         | 3         |
| Exposure time [s]                             | 0.1       | 0.1       |
| Detector pixel size [μm]                      | 18        | 18        |
| f-number                                      | 2         | 2         |
| System efficiency                             | 0.5       | 0.5       |
| Readout noise [electrons]                     | 60        | 60        |
| Airborne observation altitude [km]            | 10        | 10        |
| Spaceborne observation altitude [km]          | 600       | 600       |
| Airborne along-track [m]                      | 20        | 20        |
| Spaceborne along-track [m]                    | 701       | 701       |
| Airborne across-track [m]                     | 2.95      | 2.95      |
| Spaceborne along track [m]                    | 155       | 155       |
| Airborne ground sampling distance [m]         | 7.7       | 7.7       |
| Spaceborne ground sampling distance [m]       | 330       | 330       |

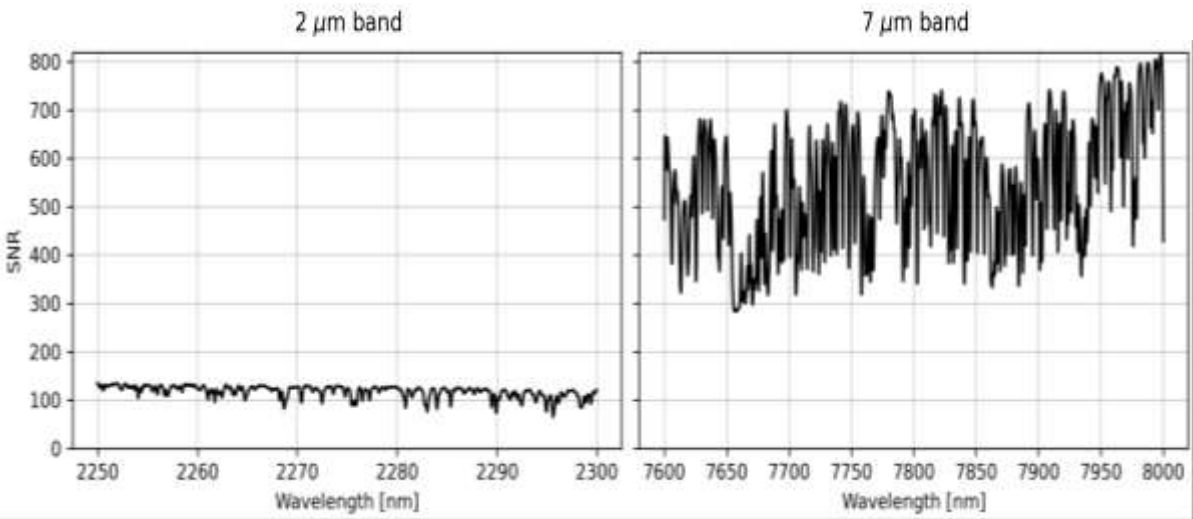
<sup>a</sup> Gaussian slit function (i.e., instrument spectral response function, ISRF) is assumed.

$$S = \frac{\pi}{4} \cdot \boxed{I} \cdot dp^2 \cdot \frac{1}{f^2} \cdot n_{\text{sample}} \cdot dt \cdot d\lambda \cdot \eta$$

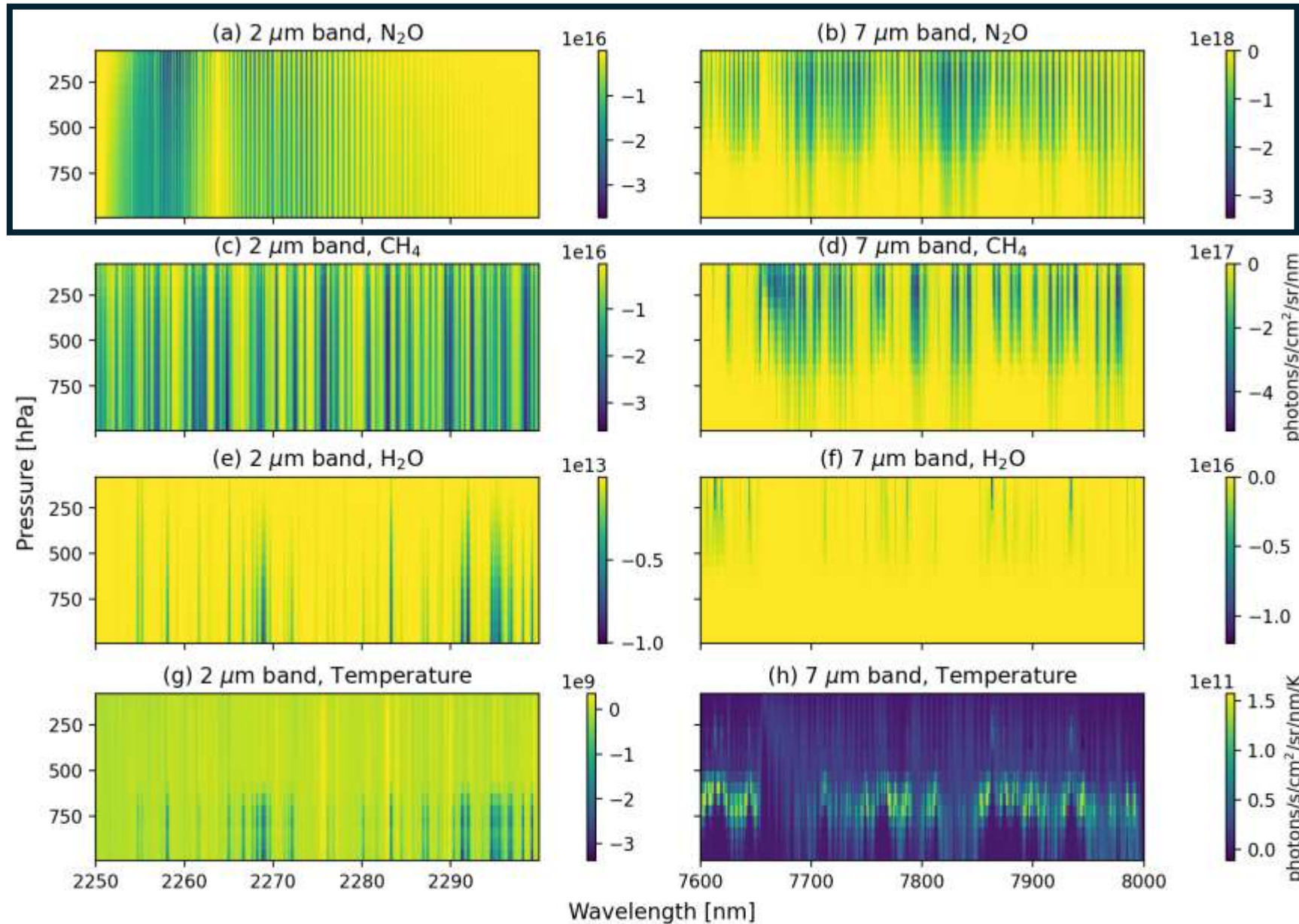
$$N = \sqrt{S + N_r^2}$$

$$\text{SNR} = \frac{S}{N}$$

SPLAT-VLIDORT radiative transfer model is used to simulate radiance (I)

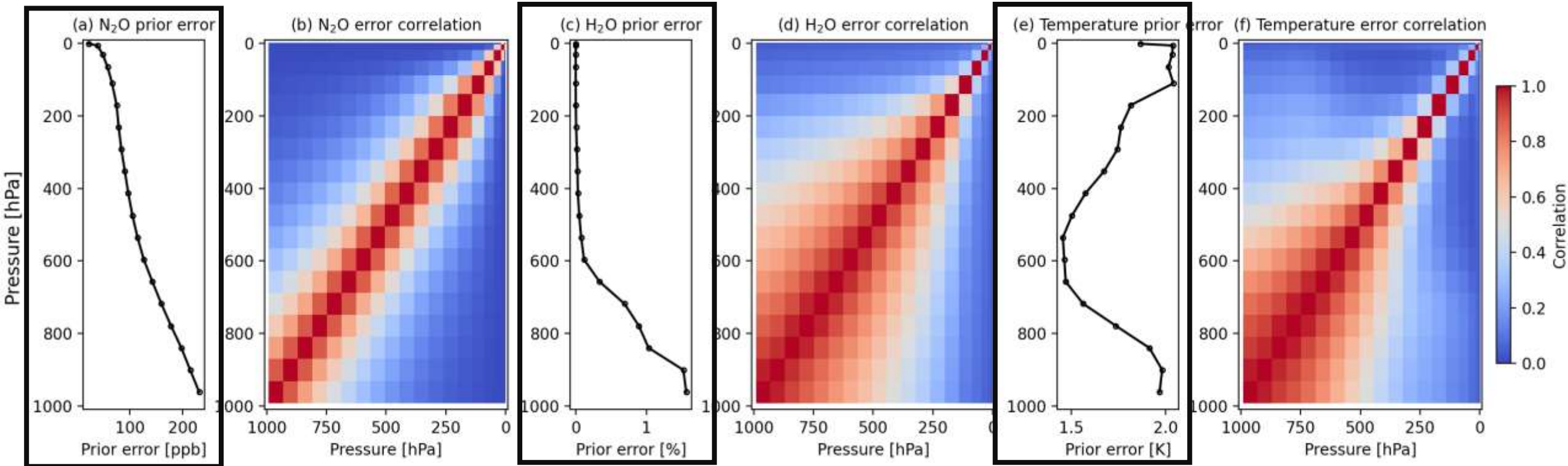


# Radiance sensitivity to state vector (Jacobians)



- State vector includes  $\text{N}_2\text{O}$ ,  $\text{CH}_4$ , water vapor, temperature profiles, and surface temperature
- 2  $\mu\text{m}$ : sensitivity to the whole column including near-surface
- 7  $\mu\text{m}$ : stronger sensitivity to  $\text{N}_2\text{O}$  as compared to the 2  $\mu\text{m}$  band, but less sensitive to near-surface layers

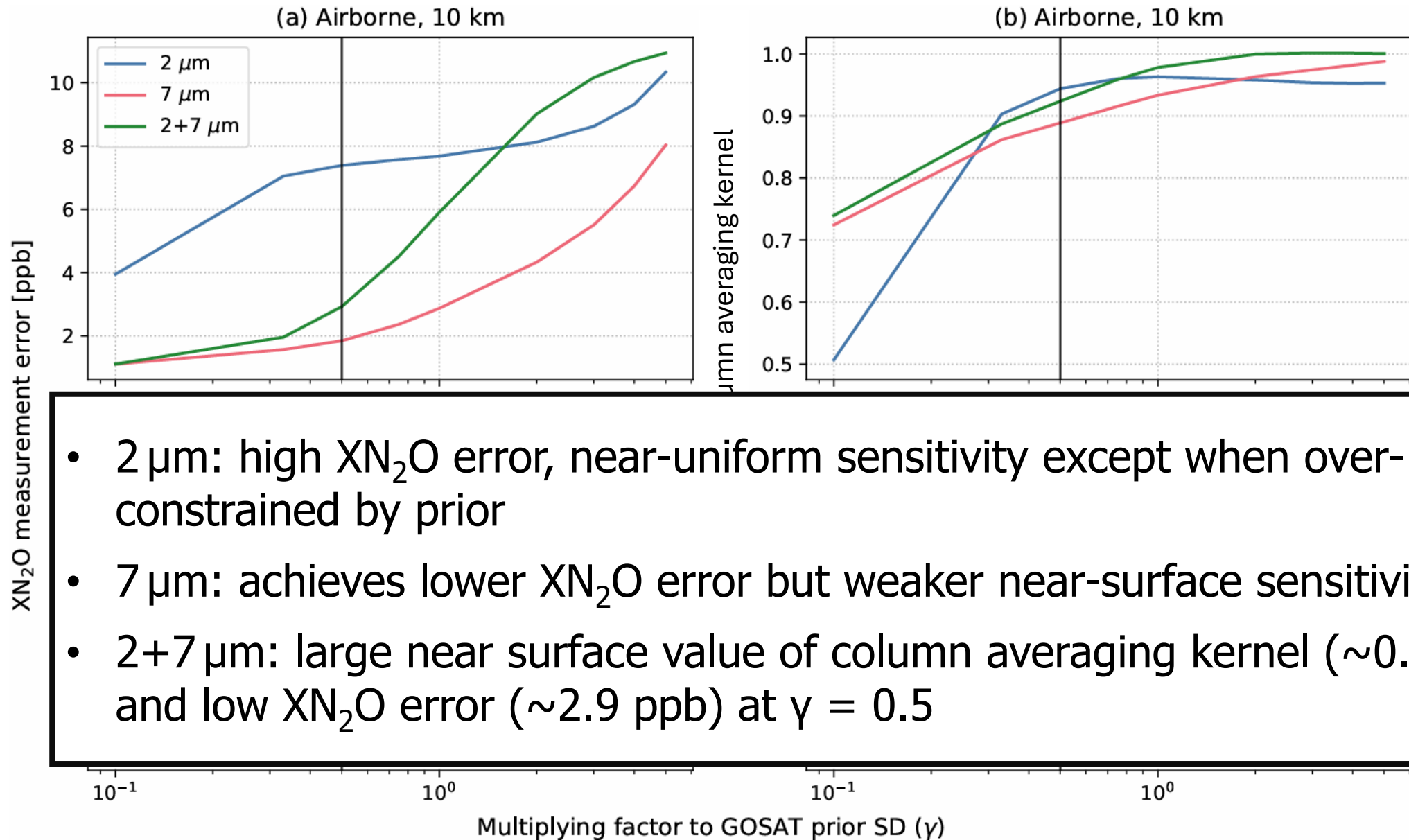
# A priori constraint



- N<sub>2</sub>O prior error is derived from University of Leicester GOSAT CH<sub>4</sub> retrieval, scaled by their background concentration ratio (330 ppb / 1900 ppb)
- H<sub>2</sub>O and temperature priors are adopted from CrIS algorithm
- Scaling factor gamma ( $\gamma$ ) is applied to N<sub>2</sub>O prior standard deviation to tune prior constraint strength

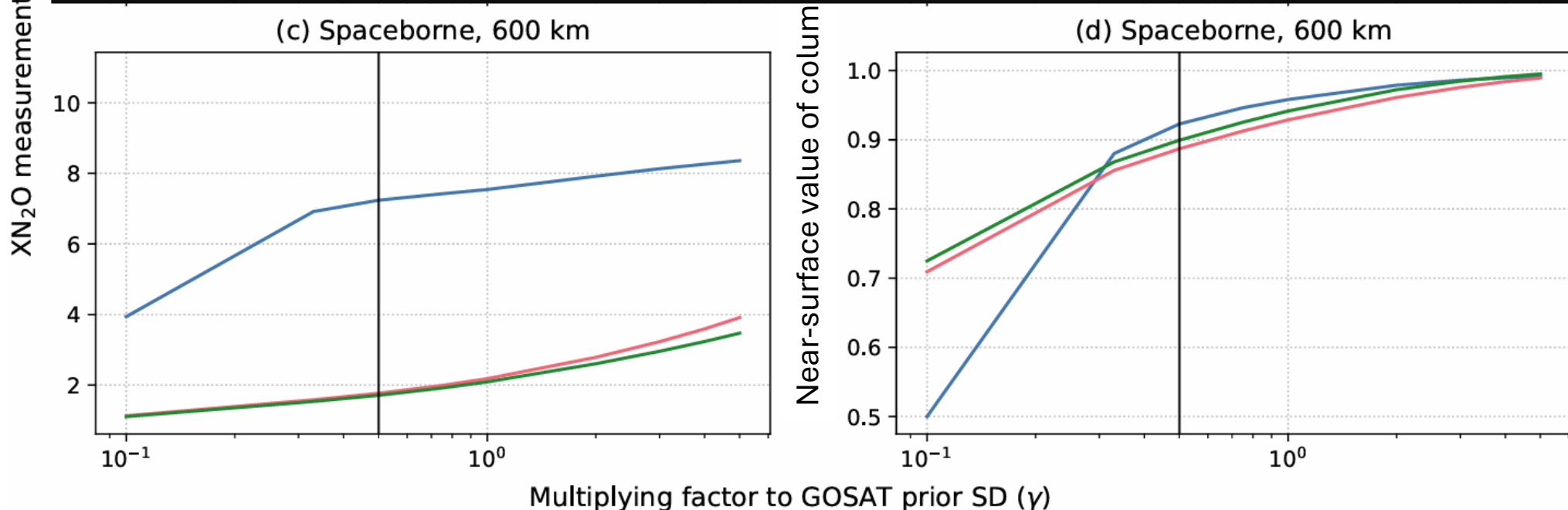


# XN<sub>2</sub>O precision



# XN<sub>2</sub>O precision

- 2  $\mu\text{m}$ : highest XN<sub>2</sub>O error across all prior strengths
- 7  $\mu\text{m}$ : comparable XN<sub>2</sub>O error with dual-band approach but weaker near-surface sensitivity
- 2+7  $\mu\text{m}$ : large near surface value of column averaging kernel ( $\sim 0.9$ ) and low XN<sub>2</sub>O error ( $\sim 1.7$  ppb) at  $\gamma = 0.5$



# How precisely can we observe $\text{XN}_2\text{O}$ ?

Using 2  $\mu\text{m}$  and 7  $\mu\text{m}$  integration and at a moderate prior strength ( $\gamma = 0.5$ ), we achieve:

- 2.9 ppb  $\text{XN}_2\text{O}$  error for airborne
- 1.7 ppb  $\text{XN}_2\text{O}$  error for spaceborne
- $\sim 0.9$  near surface value of column averaging kernel for both

**With those precisions, can we detect actual surface emissions, and at what spatial scale?**

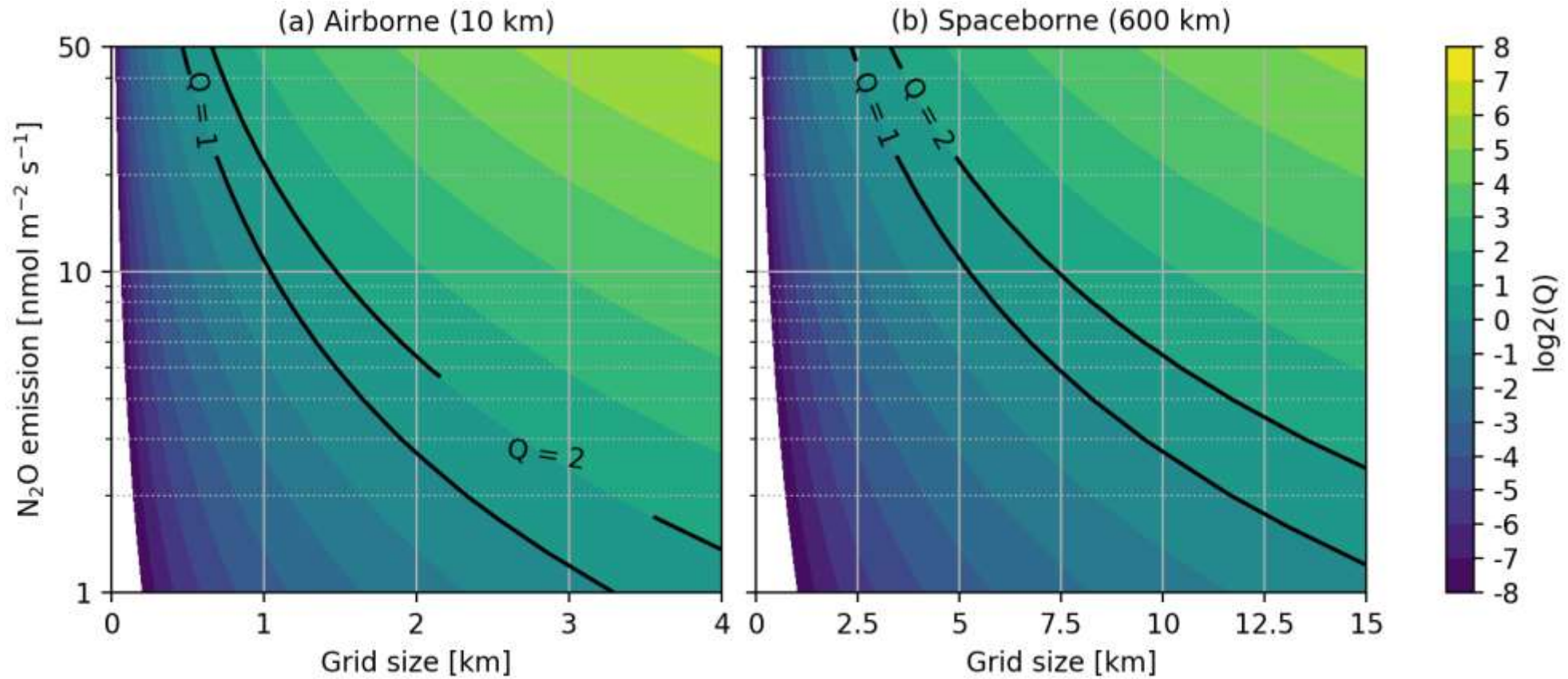
# Estimating N<sub>2</sub>O emissions using errors

$$Q = \frac{\text{Emission induced enhancement } (\Delta\Omega)}{\text{Column error } (\sigma_\Omega)} = \frac{E\Delta x^2 M_{\text{air}} g}{|u| \sigma_{\text{XN}_2\text{O}} x_0 P_{\text{air}}}$$

| Parameter                                                   | Value [unit]                                 |
|-------------------------------------------------------------|----------------------------------------------|
| Emissions (E)                                               | 1-50 [nmol m <sup>-2</sup> s <sup>-1</sup> ] |
| Airborne ground sampling distance (x <sub>0_air</sub> )     | 7.7 [m]                                      |
| Spaceborne ground sampling distance (x <sub>0_space</sub> ) | 330 [m]                                      |
| Airborne aggregation scale (Δx)                             | X <sub>0_air</sub> to 4 [km]                 |
| Spaceborne aggregation scale (Δx)                           | X <sub>0_space</sub> to 15 [km]              |
| Molar mass of dry air (M <sub>air</sub> )                   | 0.029 [kg mol <sup>-1</sup> ]                |
| Acceleration due to gravity (g)                             | 9.8 [m s <sup>-2</sup> ]                     |
| Wind speed ( u )                                            | 1.389 [m s <sup>-1</sup> ]                   |
| Airborne σ <sub>XN<sub>2</sub>O</sub>                       | 2.9 [ppb]                                    |
| Spaceborne σ <sub>XN<sub>2</sub>O</sub>                     | 1.7 [ppb]                                    |
| Pressure (P <sub>air</sub> )                                | 1e5 [Pa]                                     |



# Estimating N<sub>2</sub>O emissions using errors

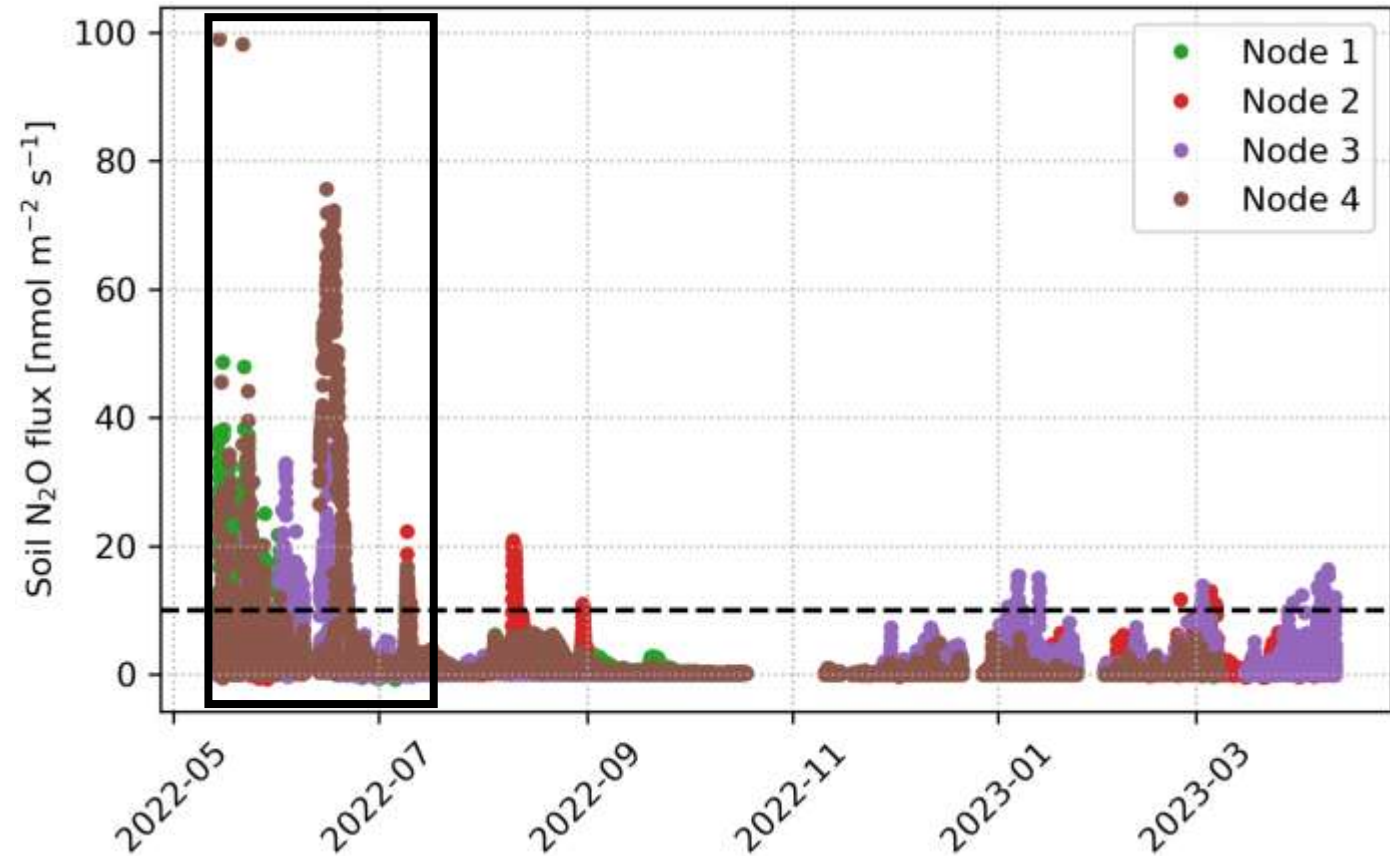


- Airborne: 10 nmol m<sup>-2</sup> s<sup>-1</sup> emission flux is observable at unit signal-to-noise ratio (Q) at ~1 km
- Spaceborne: 10 nmol m<sup>-2</sup> s<sup>-1</sup> emission flux is observable at unit signal-to-noise ratio (Q) at ~5 km

# N<sub>2</sub>O chamber data from University of Illinois



**Auto-chambers on conventionally tilled maize field in central Illinois, USA**



- 4 Nodes, 5 chambers per node
- $5 \times 10^4 \text{ m}^2$  area of the field, with nodes 50–100 m distance apart
- $10 \text{ nmol m}^{-2} \text{ s}^{-1}$  or higher N<sub>2</sub>O emission does occur in the real field especially during growing season
- N<sub>2</sub>O flux correlated for nodes that are 100 m apart



# Conclusion

- Evaluated remote sensing solutions for high-resolution mapping of  $\text{N}_2\text{O}$  flux variability in agricultural landscapes
- Integrating the shortwave and longwave bands deliver strong near-surface sensitivity ( $\text{AVK} \approx 0.9$ ) and low  $\text{XN}_2\text{O}$  error ( $\approx 2.9$  ppb airborne; 1.7 ppb spaceborne)
- Emission flux of  $10 \text{ nmol m}^{-2} \text{ s}^{-1}$  is observable at  $Q = 1$  down to  $\sim 1$  km for airborne and  $\sim 5$  km for spaceborne